# K<sup>-</sup> nucleon/multi-nucleon interaction studies by AMADEUS towards clarifying the existence of kaonic nuclear states

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The existence of Kaonic Nuclear States of K<sup>-</sup>, also called kaonic nuclear clusters, was already predicted in 1986 [1]. Since then they have been intensively debated by the scientific community, both by experimentalists and by the theoreticians. According to some theories the existence of very deeply bound states is possible while other theories are predicting much less bound states [2-4]. Therefore, in order to clarify this issue, experimental data are needed. AMADEUS goal is to do the first complete investigation of the  $\Lambda p$ ,  $\Sigma^0 p$ ,  $\Lambda d$ ,  $\Sigma^0 d$  and  $\Lambda t$  channels, searching for signals coming from the possible bound states and, in the same time, exploring intensively the rich physics of these channels [5-8]. The absorption of low momentum K<sup>-</sup>mesons ( $p_{\kappa} \sim 127$  MeV/c), produced by the DA $\Phi$ NE [1] collider, on He and C nuclear targets is investigated, with the aim to get information on the resonant transition amplitudes few MeV below the KN threshold, which represent good tests for the theoretical predictions of the low energy QCD models in the strangeness sector. The measurement of K<sup>-</sup> multiN BRs and low-momentum cross sections, are an essential tool for the investigation of the possible existence of K K<sup>-</sup> multiN bound states and for the investigation of the Kbar properties in nuclear medium.

the

neutron star with pion condensate

10<sup>6</sup> g/cm <sup>3</sup>

10<sup>14</sup> g/cm <sup>3</sup>

Hydrogen/He atmosphere

## 2. Motivation

nuclear and particle physics:

#### > astrophysics:

11]

expected

including

gravitational waves.

strange star

to

the binding of the kaon in nuclear medium may impact models on studies of in-medium modification of KN interaction for low energies in non-describing the structure of neutron stars

(Equation of State of neutron stars) [9-

be

n,p,e, µ

binaries which are

sources of

R ~ 10 km

aditional neutron sta

## 5. K<sup>-</sup> - multiN absorption and search for DBKNS $K^{-4}He \rightarrow \Lambda/\Sigma^{0} t$ (4NA)

Ecal

 $K^{-12}C \rightarrow \Lambda/\Sigma^0 p (2NA)$ 

#### perturbative QCD

solving following problems: (i) hadron masses (related to the chiral symmetry breaking), (ii) hadron interactions in nuclear medium and (iii) structure of the dense nuclear matter [2-4].

ChPT not applicable to the KN channel due to the  $\Lambda(1405)$  and the  $\Sigma(1385)$ resonances just below the KN mass threshold

Σ (1385) Λ <mark>(1405)</mark> 1500  $\sqrt{s}$  [MeV Λπ Σπ KN

**Possible solutions:** 

-Non-perturbative Coupled Channels Approach (Chiral Unitarity SU(3) Dynamics) -Phenomenological KN and NN potentials

## **3. AMADEUS experiment**

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## DAΦNE

- $\phi \rightarrow K^- K^+ (49.2\%), \approx 1000 \phi/s$
- monochromatic low momentum Kaons ≈127 Mev/c
- **back to back** K<sup>-</sup> K<sup>+</sup> topology
- small hadronic background due to







# **KLOE**

- Cilindrical DC with  $4\pi$  geometry &
- electromagnetic calormeter
- 96% acceptance
- high efficiency and resolution for charged and neutral particles
- exclusive measurement of the considered processes
  - K<sup>-</sup> absorption on light nuclei **AT REST & IN FLIGHT**

### 4. K<sup>-</sup>"n" $\rightarrow \Lambda \pi^-$ resonant vs. non-resonant

Non-resonant transition amplitude





Table 1 Branching ratios and cross sections of the K<sup>-</sup> multi-nucleon absorption processes. The statistical and systematic errors are also shown.

Process	Banching Ratio (%)	$\sigma$ (mb) @ $p$	$p_K (MeV)$	V/c)
2NA-QF $\Lambda p$	$0.25 \pm 0.02 \text{ (stat.)} ^{+0.01}_{-0.02} \text{(syst.)}$	$2.8 \pm 0.3 \text{ (stat.)} ^{+0.1}_{-0.2} \text{ (syst.)}$	0	$128\pm29$
2NA-FSI $\Lambda p$	$6.2 \pm 1.4$ (stat.) $^{+0.5}_{-0.6}$ (syst.)	$69 \pm 15 \text{ (stat.)} \pm 6 \text{ (syst.)}$	0	$128\pm29$
2NA-QF $\Sigma^0 p$	$0.35 \pm 0.09 (\text{stat.}) \stackrel{+0.13}{_{-0.06}} (\text{syst.})$	$3.9 \pm 1.0 \text{ (stat.)} ^{+1.4}_{-0.7} \text{ (syst.)}$	0	$128\pm29$
2NA-FSI $\Sigma^0 p$	$7.2 \pm 2.2$ (stat.) $^{+4.2}_{-5.4}$ (syst.)	$80 \pm 25 \text{ (stat.) } ^{+46}_{-60} \text{ (syst.)}$	0	$128\pm29$
$3NA \Lambda pn$	$1.4 \pm 0.2$ (stat.) $^{+0.1}_{-0.2}$ (syst.)	$15 \pm 2 \text{ (stat.)} \pm 2 \text{ (syst.)}$	0	$117\pm23$
3NA $\Sigma^0$ pn	$3.7 \pm 0.4$ (stat.) $^{+0.2}_{-0.4}$ (syst.)	$41 \pm 4 \text{ (stat.)} {}^{+2}_{-5} \text{ (syst.)}$	0	$117\pm23$
4NA Apnn	$0.13 \pm 0.09 (\text{stat.}) \stackrel{+0.08}{_{-0.07}} (\text{syst.})$	$\mathbf{p} \mathbf{p} (\mathbf{u} - 12, \mathbf{c}) = \mathbf{p} (\mathbf{p} - 12, \mathbf{c})$	0.00	t < <u>+0.027</u>
2NA- $\Sigma/\Lambda$ conv.	$2.1 \pm 1.2$ (stat.) $^{+0.9}_{-0.5}$ (syst.)	$BR(K^{-12}C \to \Lambda(\Sigma^0)pR) = 0.177 =$	$\pm 0.024$	$4(stat.)^{+0.027}_{-0.032}$





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