





UNIA EUROPEJSKA EUROPEJSKI FUNDUSZ ROZWOJU REGIONALNEGO



INTERNATIONAL PHD PROJECTS IN APPLIED NUCLEAR PHYSICS AND INNOVATIVE TECHNOLOGIES

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FAZIA detector system - motivation, first results and current status

Tomasz Twarog

Jagiellonian University

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Outline

Drawbacks of existing 4π arrays Further motivation First physics results Current status

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Further motivation

First physics results

Current status

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INDRA CHIMERA

 $\mathsf{INDRA} = \mathsf{Identification}$ des Noyaux et Detection a Resolution Accrue

Design:

- wide angular coverage: $\sim 90\%$ of 4π
- high granularity
- small detection thresholds of $\sim 1 MeV/A$
- ▶ high charge resolution (up to Z~50)
- isotopic identification of light charged particles

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INDRA CHIMERA

How it looks:



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INDRA CHIMERA

Characteristics:

- composed of ionisation chamber silicon -CsI(TI) telescopes
- ΔE -E and light shape analysis used to identify nuclei
- ▶ good identification of charged products up to Z~60 with a threshold of ~ 1 AMeV
- problem: isotopic identification limited to Z = 6

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INDRA CHIMERA

CHIMERA = CHarged Ion Mass and Energy Resolving Array 4π detector for charged particles devoted to the study of nuclear reactions at intermediate energies (10-100 MeV/A) and operating at Laboratori Nazionali del Sud in Catania



INDRA CHIMERA

Characteristics:

- ensemble of silicon CsI(TI) telescopes
- ► identification through ToF, ΔE-E and light shape analysis techniques
- very low thresholds for mass identification, but much higher ones for Z identification
- isotopic identification up to Z = 13 still not satisfactory

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Spectrometers:

- still unmatched nucleus identifiers
- A, Z, energy and velocity provided with high resolution
- problem: covering a very small solid angle, which makes it impossible to fully analyse many-body events, unless coupled to other detection array

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FAZIA:



Interest: isospin dependence of the competition between multifragmentation and fission channels

What can be learnt/achieved:

- discrimination of theoretical models the ratio of the two channels still unknown
- isospin effect on energy:

$$E(N,Z) = -a_V A + a_S A^{2/3} + a_{sym} \frac{(N-Z)^2}{A} + a_C \frac{Z^2}{A^{1/3}}$$

- symmetric nuclear matter extensively explored, but asymmetric (N − Z ≠ 0) largely unknown
- systems very asymmetric in neutron and proton numbers need to be explored

Motivation (continued):

- study of the de-excitation properties of isotopic chains of compound nuclei, with complete isotopic identification of the emitted nuclei and of the final residue, shall give unique information on the temperature dependence of E_{sym}
- pushing CN towards the drip lines will permit to explore decay modes that cannot be predicted by the Weisskopf theory (clustering, multifragmentation)
- at low energies, dissipative collisions induced by exotic projectiles should provide new information on the neck-instabilities, and on the achievement of isospin equilibration, both directly depending on E_{sym}

Recent results from FAZIA experiment:

- Kr beam impinging on neutron-rich and neutron-poor isotopes
- isospin diffusion between target and projectile



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- good isotopic resolution of telescope allows to investigate isotopic composition of fragments
- similarity to C and Mg isotopes (shown above)
- ▶ good agreement with ¹³⁶Xe + p
- conclusion: final fragment isospin content does not depend on the dynamics, but original neutron content is "remembered"



 $\begin{array}{c} \text{Outline} \\ \text{Drawbacks of existing } 4\pi \text{ arrays} \\ \text{Further motivation} \\ \textbf{First physics results} \\ \text{Current status} \end{array}$

- Dependence of isospin content on the phase-space region?
 - isospin diffusion seen again
 - breakup of neck-like structure formed between QP and QT
 - lighter fragments from the central part of the neck have small velocities
 - very high values of (N) /Z indicate isospin drift



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Possible improvements:

- replacing the FAZIA modules at backward angles with Single Chip Telescopes: reduction in the number of electronic channels (hence cost and complexity of the apparatus), as one Si detector can be used as a photodiode for the scintillator light readout
- even more advantages: less-crowded FEE, lower power dissipation
- SCTs already tested, not as good as regular telescopes for low-energetic fragments with Z > 2

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Improvements (continued):

- pulse shape analysis used for identification
- threshold significantly reduced



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- FAZIA project presently in the phase of building a demonstrator = an array of 192 telescopes
- final electronics and mechanical solutions are adopted
- FEE located inside the vacuum chamber to minimize detector-preamplifier-digitizer distances
- demonstrator expected to be ready in 2014

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Thank you for your attention

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